

# Earth Tomography with KM3NeT-ORCA

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#### ABSTRACT

KM3NeT-ORCA is a future water-Cherenkov neutrino detector of for studying the oscillations of atmospheric neutrinos, with the objective of measuring the neutrino mass ordering. ORCA deployed in the Mediterranean Sea, offshore Toulon, at a c 2450m. It will be able to reconstruct interacting neutrinos of all with a large range of baselines and energies, and an effective several megatons.

Neutrinos crossing the Earth undergo matter effects, modify pattern of their flavour oscillations. The study of the angular and distribution of neutrino events in ORCA can therefore tomographic information on the Earth interior with an indep technique, complementary to the standard geophysics methods

We study here the sensitivity to the electron density for two layers of the Earth: mantle and outer core. A full Mont simulation of ORCA, event reconstruction and flavour classification used to model the detector response. A simple statistical an then performed to derive confidence intervals for the measure

	DETECTION PRINCIPLE			EARTH MODEL BASIS		
designed primary	<ul> <li>Cherenkov light produced by secondary particles emerging from neutrino interactions is detected by multi-PMT Digital Optical Modules (DOM). The DOMs are attached to vertical strings anchored to the seabed.</li> <li>9m vertical spacing of DOMs</li> <li>20m average horizontal spacing of lines</li> </ul>			Proportionality factor between electron density and matter density depends on chemical & isotopic composition: $\frac{Z}{A} = \frac{\sum x_i Z_i}{\sum x_i A_i} = \sum w_i \frac{Z_i}{A_i} \implies n_e = \frac{N_A}{m_n} \times \frac{Z}{A} \times \rho_{\text{matter}}$		
will be depth of flavours						
mass of	<ul> <li>115 lines × 18 DOMs × 31 PMTs = 5.7 Mton water instrumented</li> </ul>			→ 3 chemically distinct layers → Z/A uniform in each layer		
ying the d energy provide pendent		<b>Track</b> signature: • $\nu_{\mu}, \overline{\nu_{\mu}}$ CC events • $\nu_{\tau}, \overline{\nu_{\tau}}$ CC with $\tau \rightarrow \mu$ • $\mu_{\text{atm}}$ contamination		Mantle: $R_{ext} \approx 6300 \text{ km}$ pyrolite (rock model) Z/A = 0.496		
s. distinct te Carlo cation, is		<b>Cascade</b> signature: • $\nu_e, \overline{\nu_e}$ CC events • $\nu_{\tau}, \overline{\nu_{\tau}}$ CC, $\tau \rightarrow had./e$		Outer core: $R_{ext} \approx 3480 \text{ km}$ pure Fe (+ 5% Ni) Z/A = 0.4656 (0.4661)		
nalysis is ement of		• NC events No $\nu/\overline{\nu}$ event-by-event		Inner core: $R_{ext} \approx 1220 \text{ km}$ pure Fe7/A = 0.4656		

the electron density in these two layers.



### $\nu + \overline{\nu}$ SIGNAL – INTERACTING EVENTS WITH NO DETECTOR EFFECTS – ORCA 10 YEARS (570 MTON.YRS)





d Iχ²l



- > Bottom: 1-D projection of absolute value  $|\chi^2(bin i)|$
- > Channel-specific ORCA effective mass superimposed

 $\rightarrow$  theoretical signal is more pronounced in muon channel (*i.e.* Tracks)

 $\rightarrow$  outer core shows higher asymmetry, but concentrated in fine-grained pattern at low energy: challenging measurement

 $E_{\text{lepton}}$ 

 $E_{\nu}$ 

- Cross-sections: GENIE event generator [4]
- Earth Model basis = 42-steps PREM model [2]

Ingredients for calculating rates of interacting events:

- Oscillation Probabilities: OscProb calculator, J. Coelho [6]
- Oscillation parameters taken from NuFit website [5],  $\delta_{CP}=0$
- Normal mass ordering assumed

SENSITIVITY EVALUATION

• Flux data tables: *Honda et. al* [3]

Exposure: ORCA instrumented vol. × 10 years = 570 Mton.yrs

#### **DETECTOR RESPONSE**

- Response in neutrino energy, zenith angle ( $\cos\theta_7 < 0$  for upgoing), and inelasticity  $Y_{\text{Bjorken}} = 1 1$
- Dedicated reconstruction algorithms for Track and Cascade channels [1]
- Event classification and atmospheric muon rejection handled by Random Decision Forest technique



Reconstruction & Classification

• Reco  $(E, \cos \theta_z, Y_{\rm Bj})$ • Event classification ("flavour ID"):  $\rightarrow$  Track / Cascade /  $\mu_{\rm atm}$ 

- Full sample of Monte Carlo events is used to fill a binned response matrix = (6+1)-dim sparse histogram
- Takes into account all efficiencies, resolutions and misidentification probabilites, with all correlations
- Validity of full Monte-Carlo approach has been checked for simple 2-D analysis presented here
- → <u>Both channels</u> contribute significantly to the sensitivity

#### Reconstructed signal in Track and Cascade channels – mantle only – ORCA 10 years 2D $\chi^2$ map (E,cos $\theta$ ) - Tracks $2D \chi^2 map (E, cos\theta)$ - Cascades Mantle composition model A: Z/A = 0.496 model B: Z/A = 0.521 (+5%) Mantle composition model A: Z/A = 0.496 model B: Z/A = 0.521 (+5%) $\chi^2 = 0.76$ $\chi^2 = 1.02$ 0.06 0.025 -0.5

## $\Delta \chi^2 = -2 \log \left[ \frac{L(x|B)}{L(x|A)} \right]$

- x = observed data A,B = model hypotheses
- Simplified method: use « Asimov dataset » instead of pseudo-experiments, then

$$\Delta \chi^2 = \sum_{\substack{\text{Tracks,}\\\text{Cascades}}} \sum_{\substack{\text{bins } E\\\text{bins } \cos \theta_z}} 2 \left[ n_A - n_B + \log \left( \frac{n_B}{n_A} \right) \right]$$

- Sum on 2D bins: reconstructed Bjorken-Y information not included (results under preparation)
- No fitted systematics / oscillation parameters
- Layers considered separately: no simultaneous measurement

#### Confidence level to reject base Z/A hypothesis in outer core and mantle







5 6 7 8 9 10

True E [GeV]

$ \begin{bmatrix} 1 & 1 & 1 & 1 \\ 0 & 1 & 1 \\ 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0$		85       80       75         70       65       60         60       0uter core       NH - tracks only         1H - tracks only       1H - tracks only         50       .36       0.38       0.4       0.42       0.44       0.46       0.48       0.5       0.52       0.54       0.56	S S S S S S S S S S S S S S		
REFERENCES	SUMMARY AND PROSPECTS				
<ol> <li>S. Adrian-Martinez et al. (the KM3NeT collaboration), <i>Letter of Intent for KM3NeT 2.0</i>, J. Phys. G: Nucl. Part. Phys. 43 (2016) 084001.</li> <li>A. M. Dziewonski and D. L. Anderson, <i>Preliminary reference Earth model</i>. Phys. Earth Planet. Inter. 25 , 297 (1981).</li> <li>M. Honda et al., <i>Atmospheric neutrino flux calculation using the NRLMSISE00 atmospheric model</i>. Phys. Rev. D92, 023004 (2015).</li> <li>C. Andreopoulos et al., <i>The GENIE neutrino Monte Carlo generator</i>. Nucl. Inst. Meth. A614 , 87 (2010).</li> <li>M.C. Gonzalez-Garcia et al., <i>Updated fit to three neutrino mixing: status of leptonic CP violation</i>. JHEP 11 (2014) 052. Website: www.nu-fit.org</li> <li>J. Coelho, <i>OscProb neutrino oscillation calculator</i>. www.apc.univ-paris7.fr/Downloads/antares/Joao/OscProb_v2.0.1.tar.gz</li> </ol>	We demonstrate the feasibility tomography with KM3NeT-ORCA the Earth mantle to an accuracy significant improvement is realise Including the reconstructed inel neutrinos from antineutrinos or reconstructions is under prepa simultaneous measurements of s	easibility of measuring the electron density in the inner layers of the Earth using neutrino oscillation eT-ORCA. After 10 years of operation, the proposed ORCA detector can measure the electron density in accuracy of 4% at 1σ confidence level. In the outer core, this performance is reduced to 9% precision. A is realised by combining the Tracks and Cascades channels. cted inelasticity in the analysis has the potential to increase the sensitivity, by allowing to distinguish trinos on a statistical basis. A full analysis quantifying the same performance with improved Bjorken-Y er preparation. Further studies will also include systematic uncertainties and a sensitivity study for nents of several layers.			